

# Li-containing scintillating bolometers for low background physics

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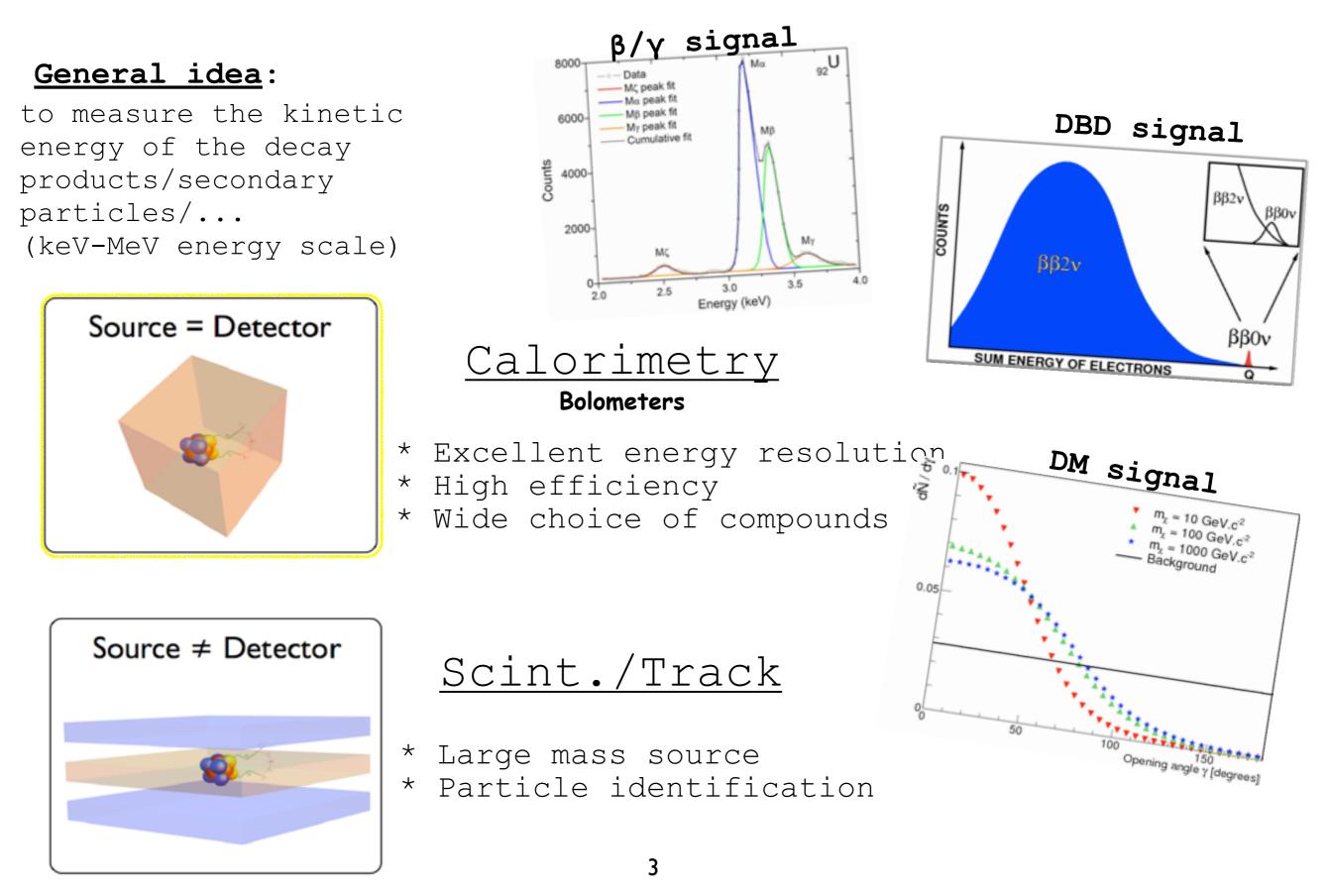
17<sup>th</sup>-20<sup>th</sup> September 2013 Kyiv,Ukraine

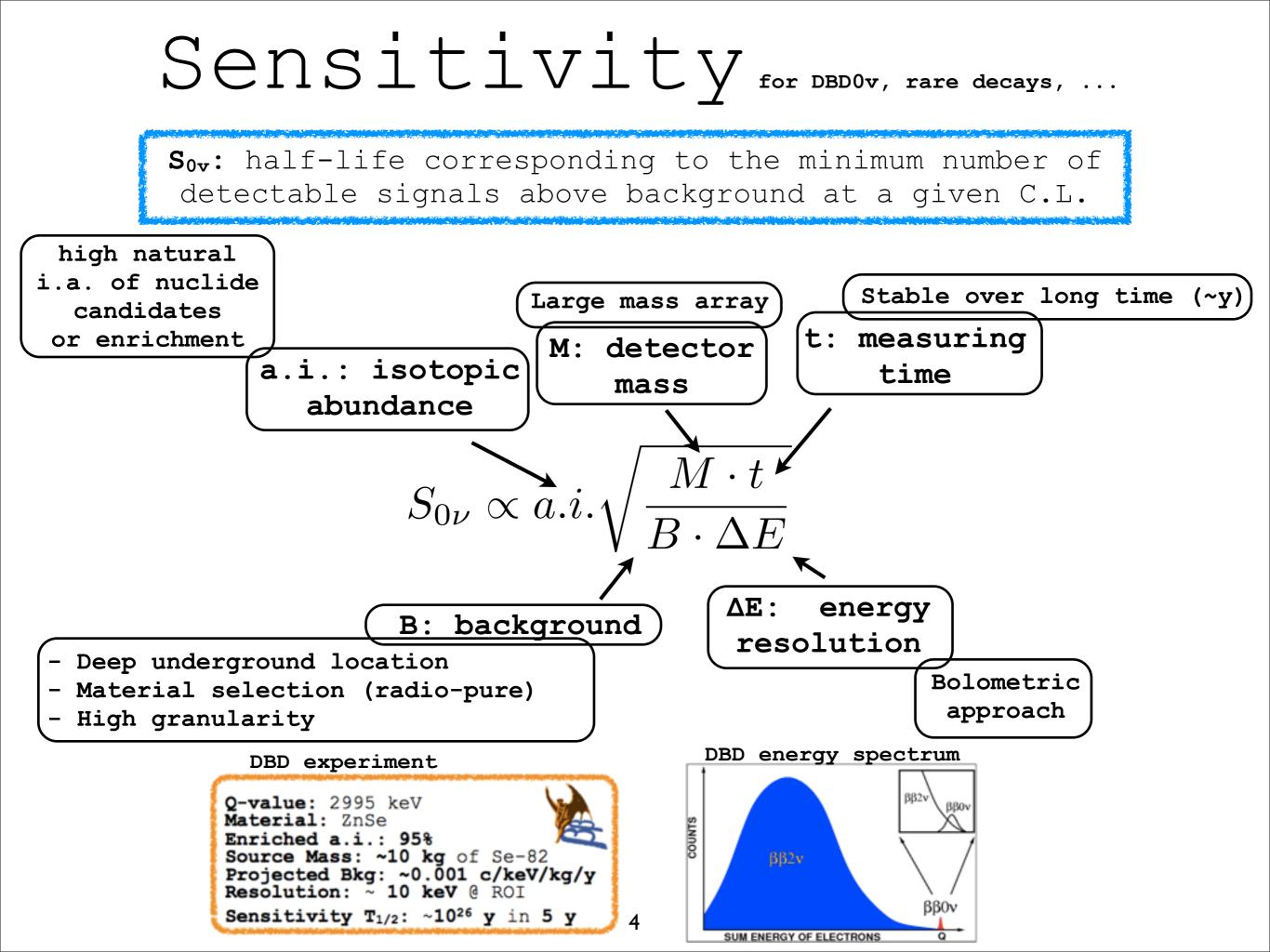


#### OUTLINE

- Rare event physics sensitivity
- The bolometric technique
  - pros and cons
- Scintillating bolometers
  - LMO & LEBO
- Conclusions

# Detection principle



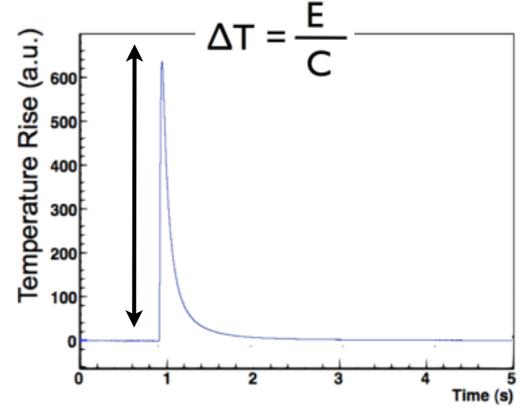


## The bolometric technique

Almost all the deposited energy is converted into phonons which induce a

measurable temperature rise

The heat capacity of the crystal must be very small (-> low Temperature ~10 mK)



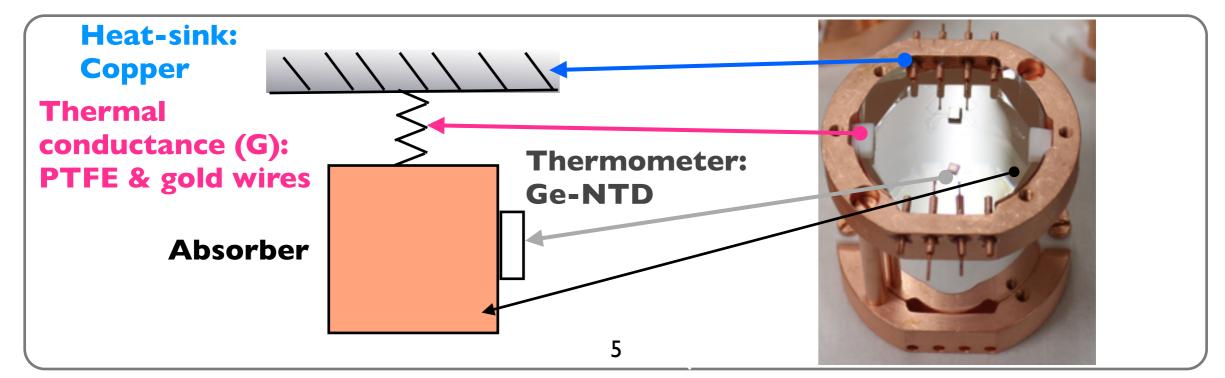
fully-active detector

#### Absorber

- M ~ 0.45 kg
- $C \sim 10^{-10} J/K$
- $-\Delta T/\Delta E \sim 500 \mu K/MeV$

#### <u>Sensor</u>

- $-R = R_0 \exp[(T_0/T)^{1/2}]$
- R  $\sim$  100 MQ
- $\Delta R/\Delta E \sim 3 M\Omega/MeV$



#### The underground facility



Laboratori Nazionali del Gran Sasso INFN, Italy

#### Experimental location:

- Average depth ~ 3650 m w.e.
- Muon flux ~ 2.6×10<sup>-8</sup>  $\mu/s/cm^2$
- Neutrons < 10 MeV:  $4*10^{-6}$  n/s/cm<sup>2</sup>
- Gamma < 3 MeV: 0.73  $\gamma/s/cm^2$

# Bkg sources in bolometric experiments

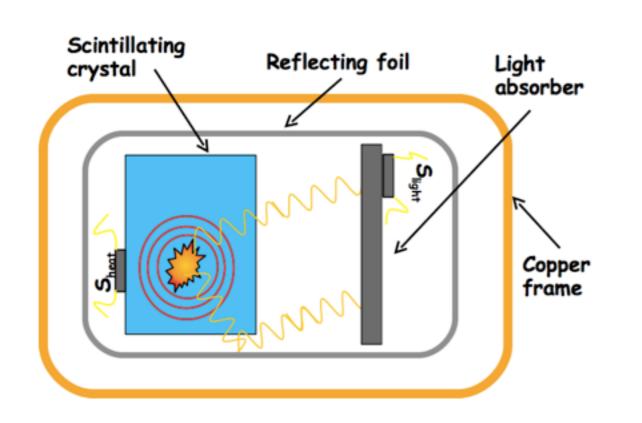
Since bolometers are **fully-active detectors** and are sensitive to all radiation types, various sources can limit the experimental sensitivity

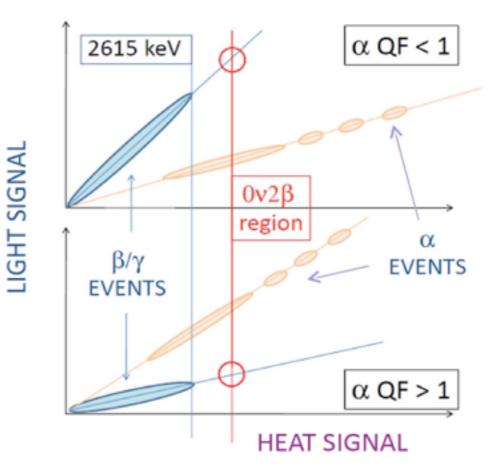
Neutrons =>	<ul> <li>neutron activation: (n,γ) reactions</li> <li>* appropriate shields are needed</li> </ul>
Muons =>	- energy deposit in the ROI * underground installation & granularity & veto
$\beta/\gamma s =>$	- natural radioactivity ( $^{238}$ U & $^{232}$ Th) * material selection
degraded $\alpha s =>$	<ul> <li>αs coming out from detector surfaces</li> <li>* surface cleaning and particle discrimination</li> </ul>

#### Scintillating bolometers

When a **bolometer is an efficient scintillator** at low temperature, a small but significant fraction of the <u>deposited</u> <u>energy is converted into scintillation photons</u> while the remaining dominant part is detected through the heat channel.

The <u>simultaneous read-out</u> of **light** and **thermal** signals allows to discriminate the  $\alpha$  background thanks to the scintillation yield different from  $\beta$  particles.

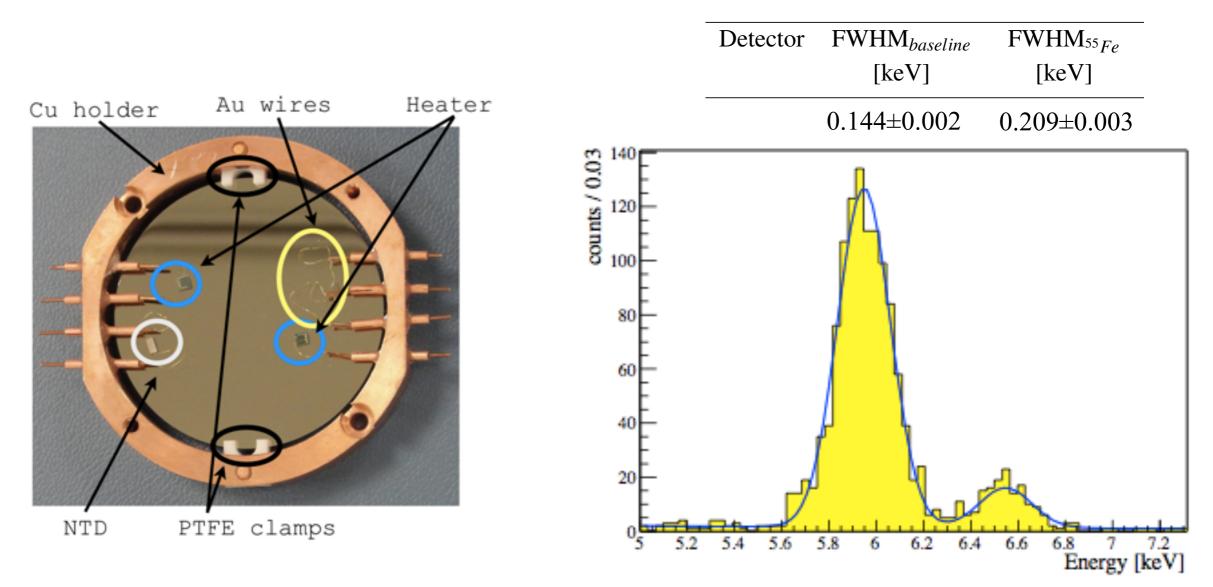




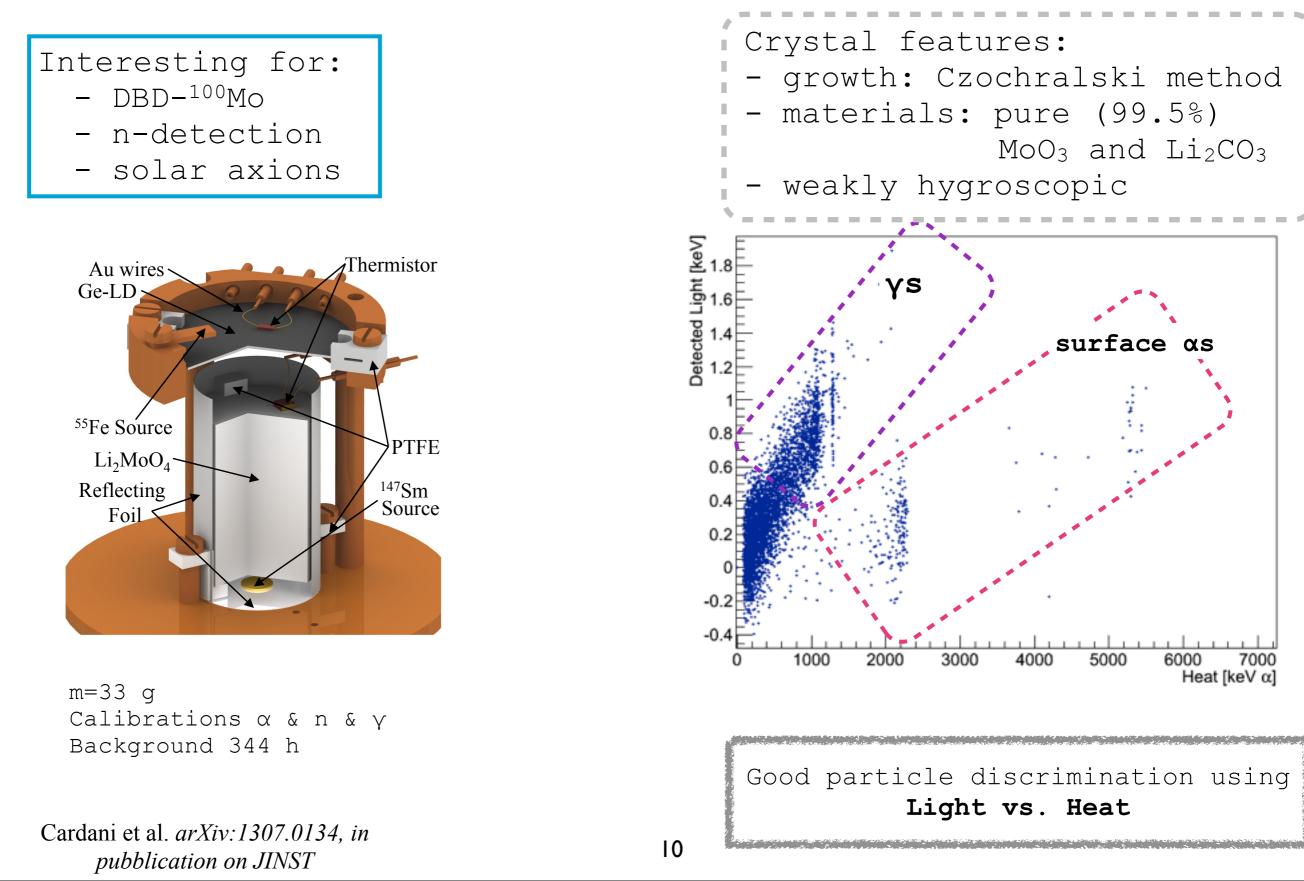
QF: is defined as the ratio of the signal amplitudes induced by an  $\alpha$  and an  $\beta/\gamma$  of the same energy.

#### Bolometric LD

- HP-Ge disk (3-5 cm diameter, 0.1-1 mm thick)
- SiO<sub>2</sub> coating for darkening the surface => reduce light reflections
- Calibration with  $^{55}\mathrm{Fe}$  X-rays @ 5.9 keV and 6.5 keV
  - Energy resolution: ~100 eV
  - Energy threshold: ~100 eV



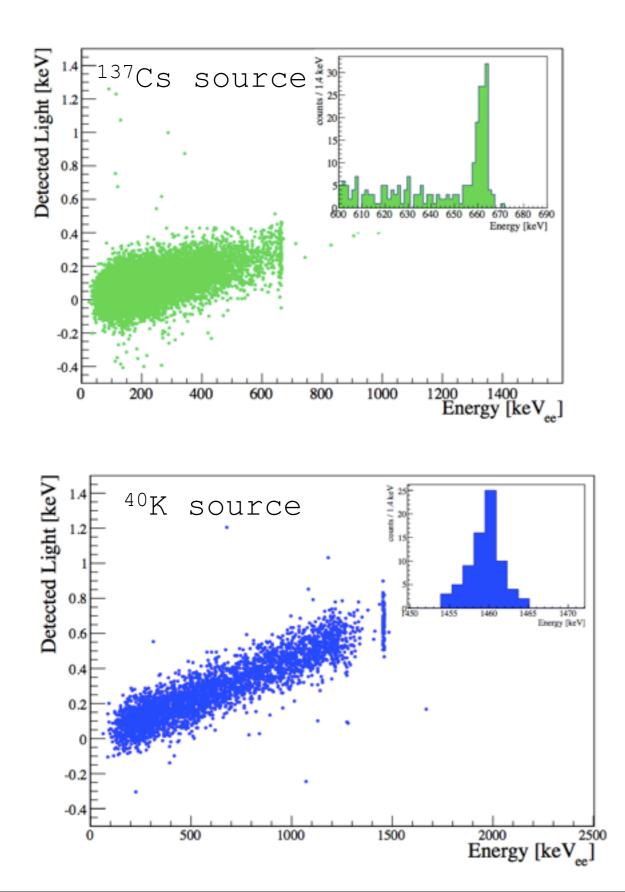
#### $Li_2MOO_4$



 $Li_2MoO_4$  with  $\gamma$ -source

П

Small crystal => no calibration at 2615 keV



FWHM: 3.7±0.8 keV @ 661 keV

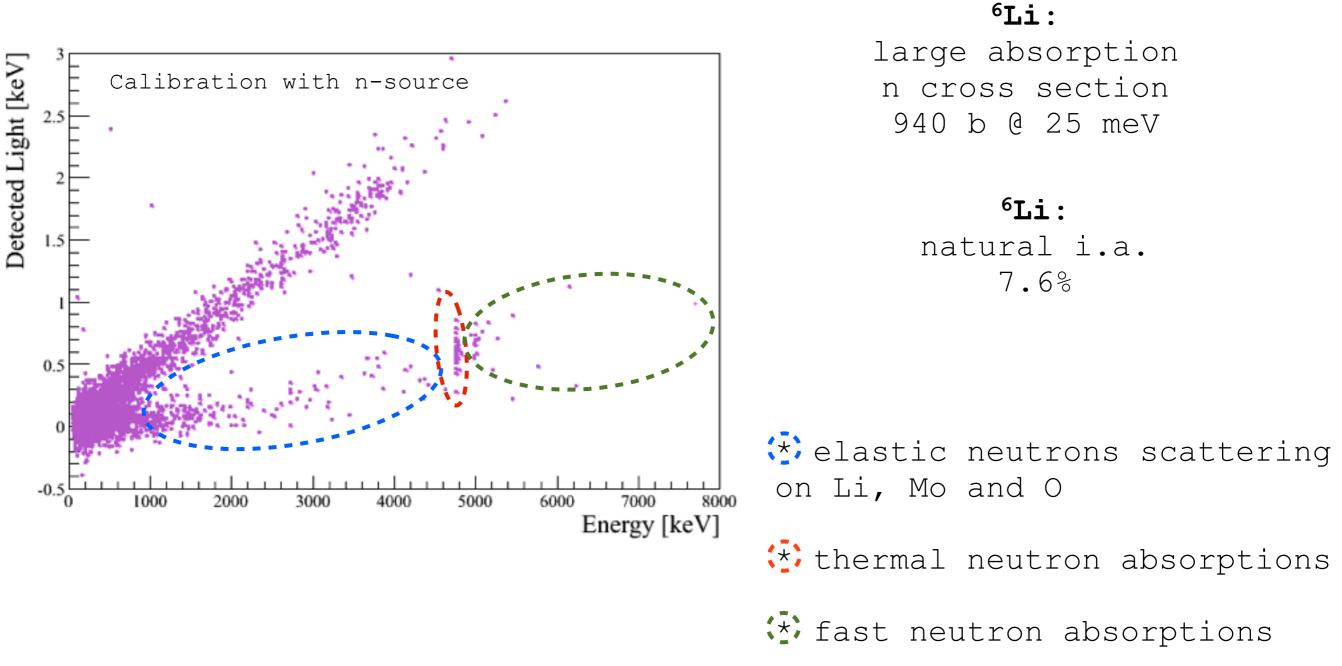
 $LY_{\beta/\gamma} = 0.433 \pm 0.012 \text{ keV/MeV}$ 

FWHM: 4.7±1.1 keV @ 1460 keV

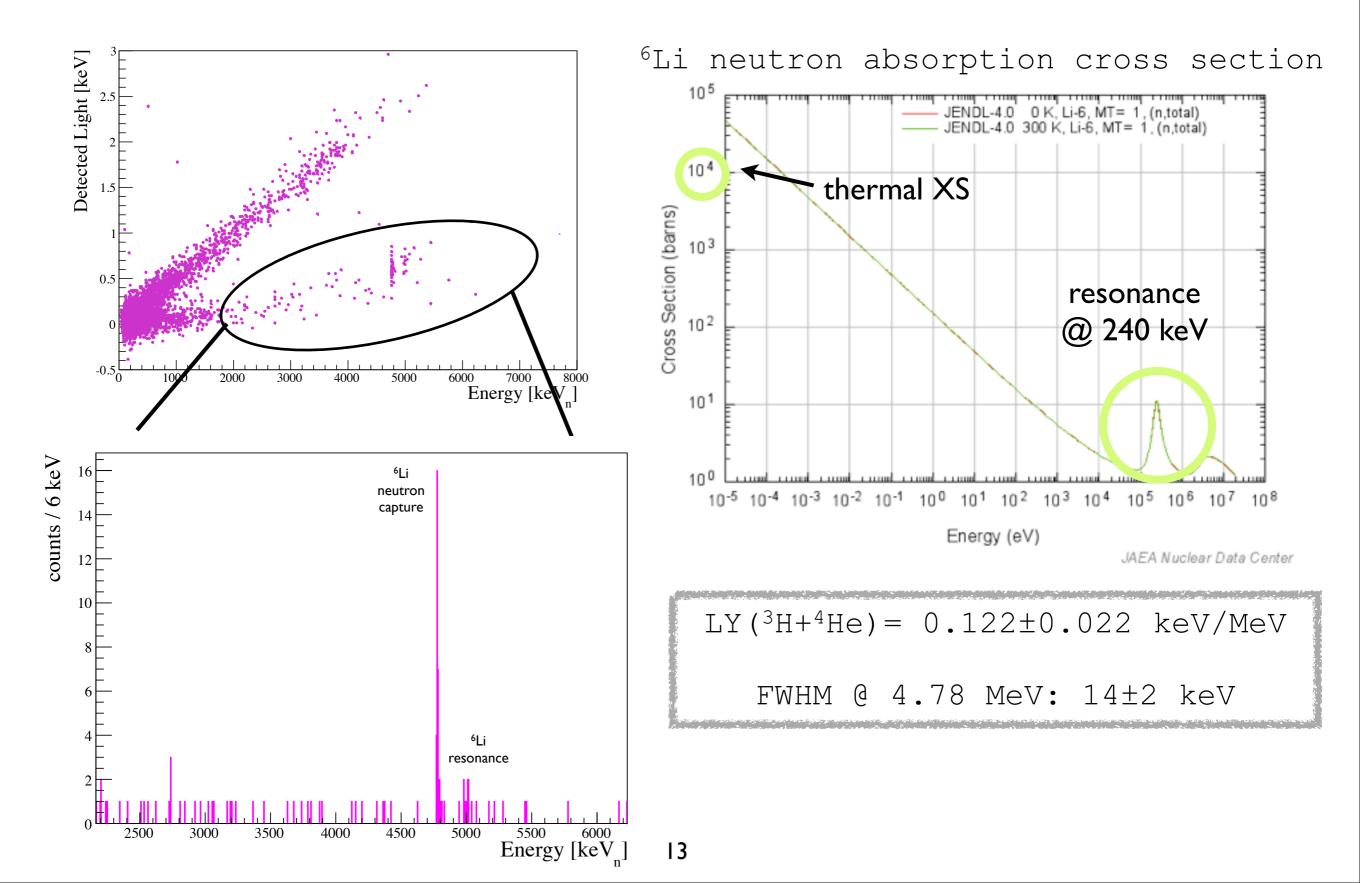
#### Li<sub>2</sub>MoO<sub>4</sub> with AmBe-source

Easy (fast) neutron tagging:

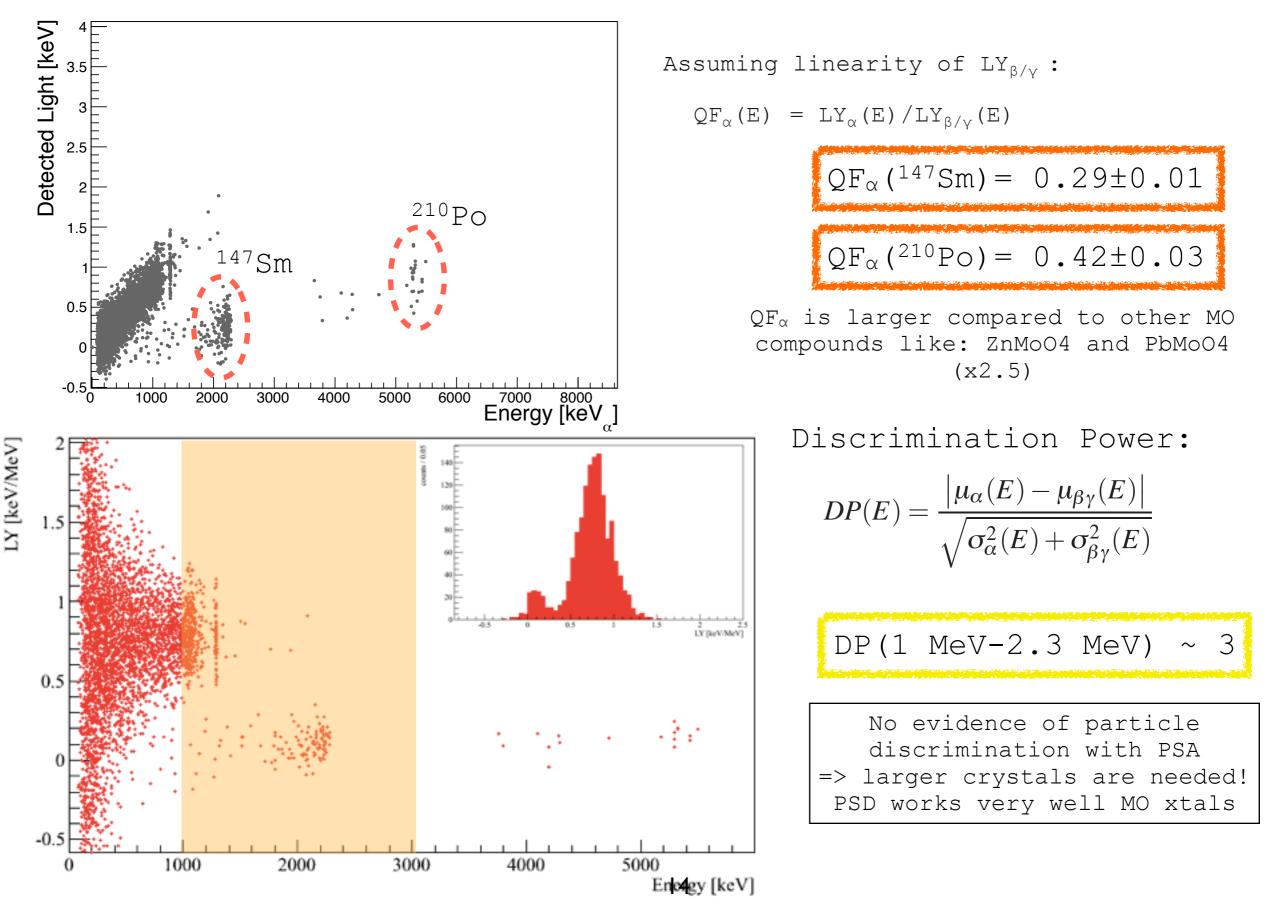
#### $^{6}\text{Li}$ + $^{1}n$ $\rightarrow$ $^{3}\text{H}$ + $^{4}\text{He}$ + 4.78 MeV

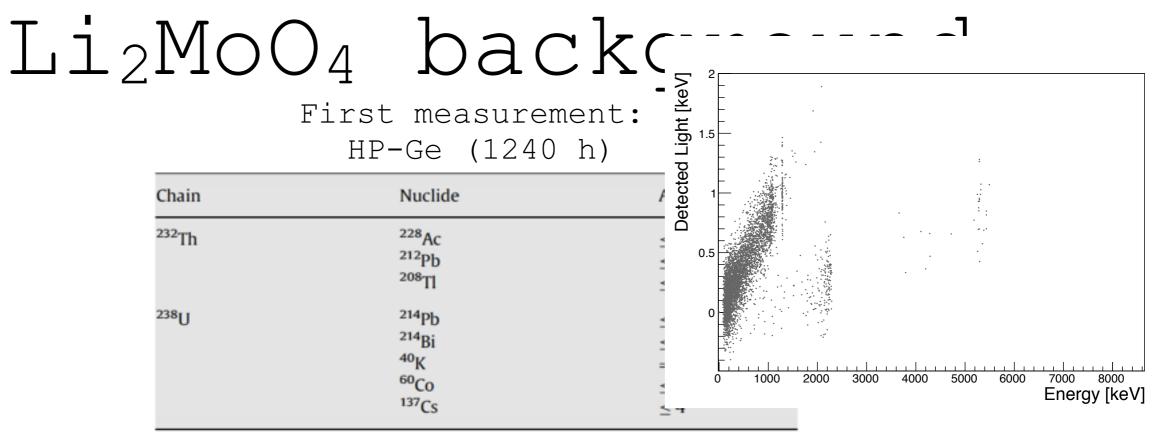


#### Li<sub>2</sub>MoO<sub>4</sub> with AmBe-source



#### $Li_2MoO_4$ with $\alpha$ -source



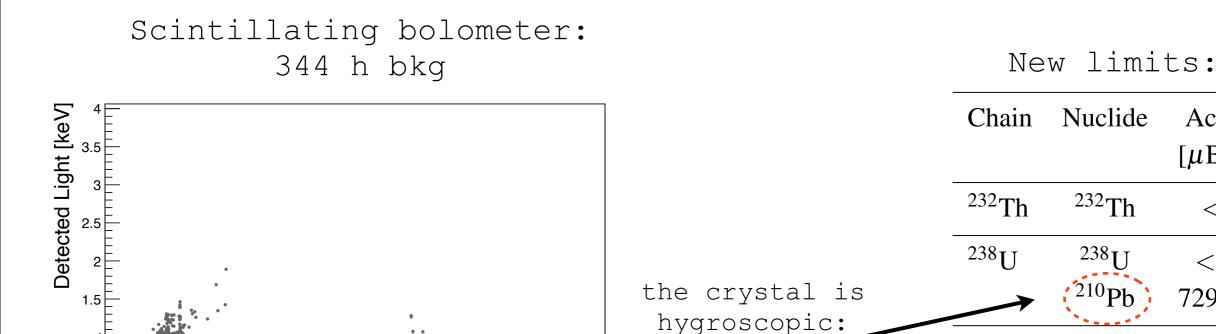


contaminated

in <sup>210</sup>Pb (<sup>210</sup>Po)

15

O.P. Barinova et al., NIM A 607 (2009) 573



7000

8000

Energy [keV ]

0.5

-0.5 **–** 0

1000

2000

3000

4000

5000

6000

Cardani et al. *arXiv:1307.0134, submitted to JINST* 

Activity

 $[\mu Bq/kg]$ 

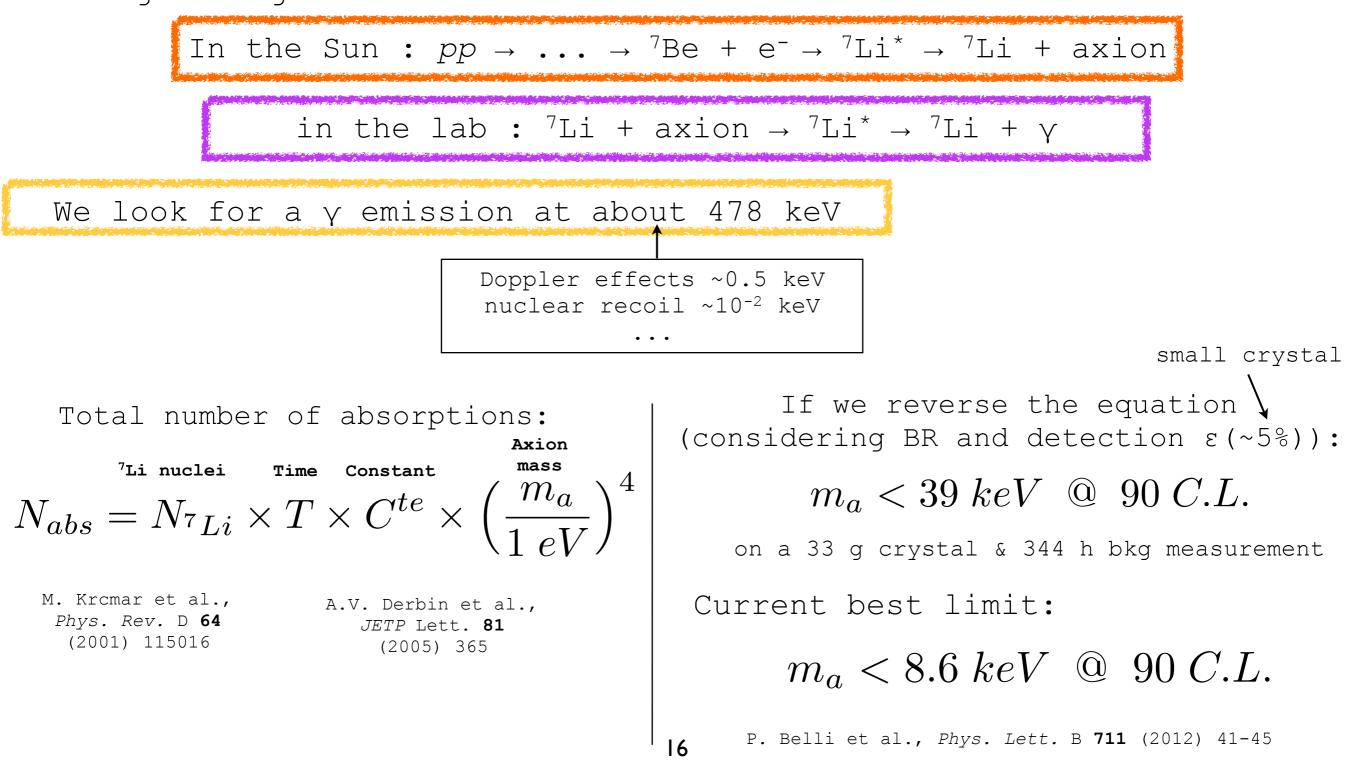
< 94

< 107

 $729 \pm 160$ 

#### Solar axions search

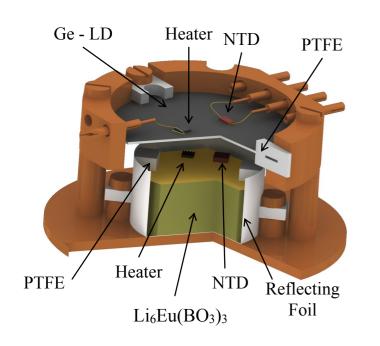
Detection of  $^{7}$ Li solar axions by means of resonant absorption on analogue targets in the labs.



## Li<sub>6</sub>Eu (BO<sub>3</sub>)<sub>3</sub>

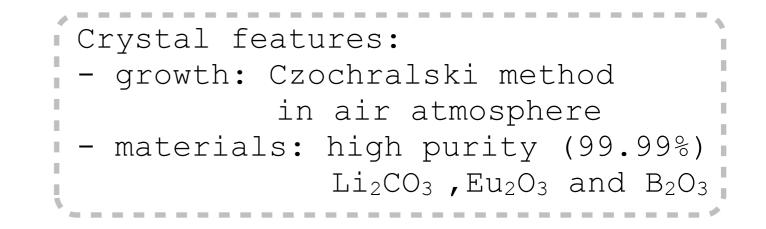
Interesting for:

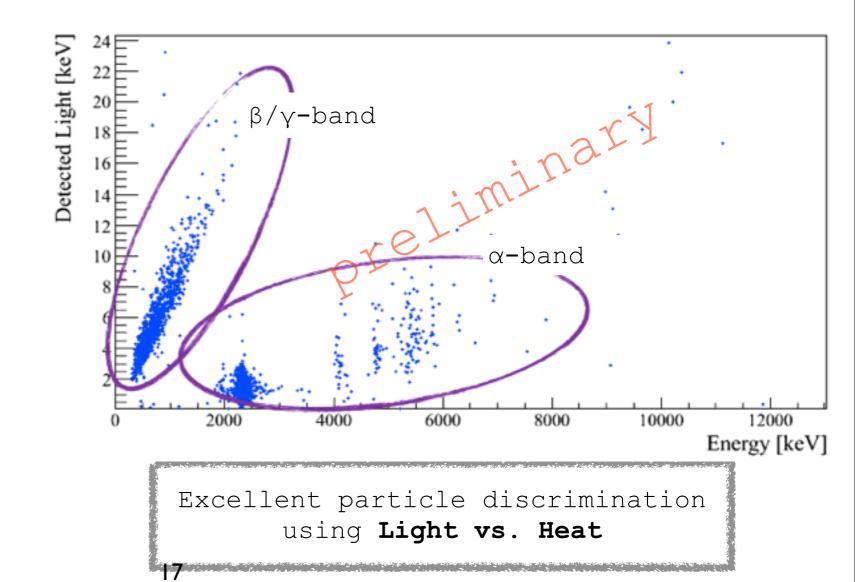
- Eu-151  $\alpha$  decay
- n-detection
- solar axions



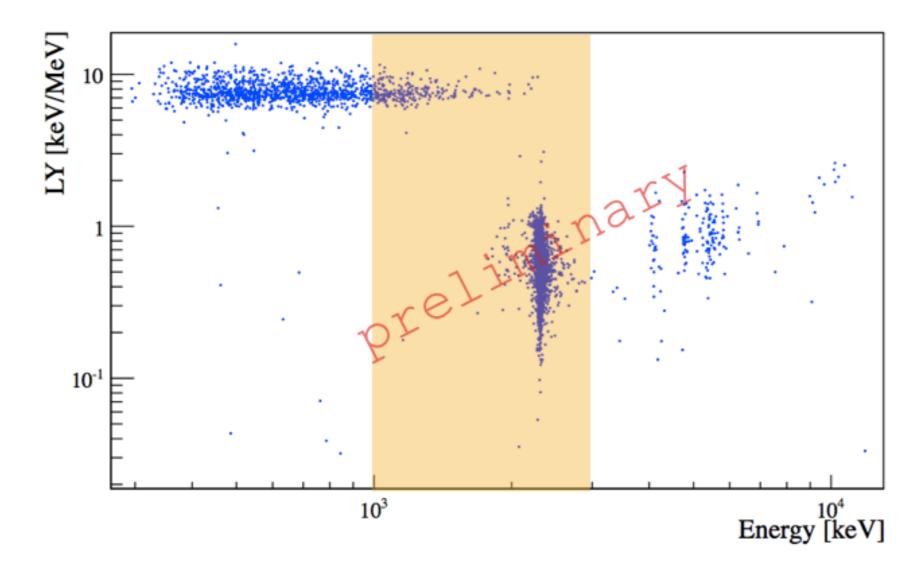
- m=6.15 g
- Calibrations  $\boldsymbol{\gamma}$
- Background

First bolometric test with 5x5x5 mm<sup>3</sup> crystal in: 2012 J. Phys.: Conf. Ser. 375 012025





#### Li<sub>6</sub>Eu (BO<sub>3</sub>)<sub>3</sub> Light Yield



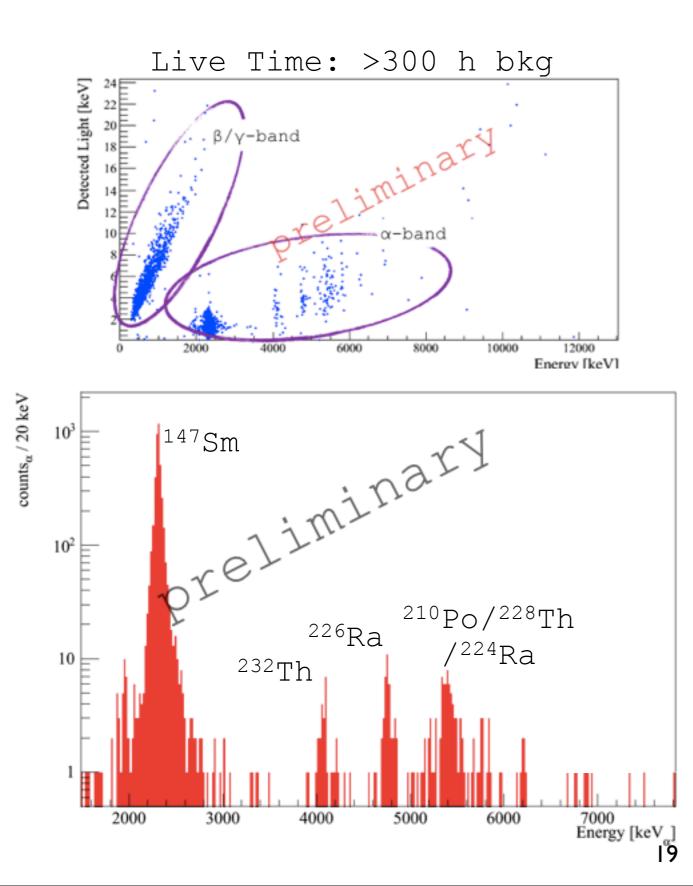
Assuming linearity of  $\mathrm{LY}_{\beta/\gamma}$  :

 $LY_{\beta/\gamma} = 7.38 \pm 0.02 \text{ keV/MeV}$ 

 $QF_{\alpha}(E) = LY_{\alpha}(E) / LY_{\beta/\gamma}(E)$ 

 $QF_{\alpha}(^{147}Sm) = 0.54\pm0.01$  $QF_{\alpha}(^{210}Po) = 0.84\pm0.05$ 

#### Li<sub>6</sub>Eu (BO<sub>3</sub>)<sub>3</sub> background



Interna	l contam	inati	ons:
Chain	Nuclide	Ac	tivity
<sup>232</sup> Th	<sup>232</sup> Th	3.5	mBq/kg
238U	<sup>238</sup> U <sup>226</sup> Ra <sup>210</sup> Po	2.9	mBq/kg mBq/kg mBq/kg
	<sup>147</sup> Sm	4.5	mBq/kg

First evaluation of intrinsic radiopurity level in 2.7 g LEBO crystal: NIM A 572 (2007) 734-738

Radioactive contaminations in Li6Eu(BO3)3 crystal

Chain	Nuclide	Activity (Bq/kg)
<sup>232</sup> Th	<sup>228</sup> Ac	< 0.20
	<sup>212</sup> Pb	< 0.25
	<sup>208</sup> Tl	< 0.13
<sup>238</sup> U	<sup>214</sup> Pb	< 0.17
	<sup>214</sup> Bi	< 0.07
	40K	<1.5
	<sup>60</sup> Co	< 0.026
	137Cs	< 0.081
	<sup>207</sup> Bi	< 0.009
	<sup>152</sup> Eu	= 0.949(48)
	<sup>154</sup> Eu	= 0.212(35)

Limits are given at 90% C.L.

## <sup>151</sup>Eu in Li<sub>6</sub>Eu (BO<sub>3</sub>)<sub>3</sub>

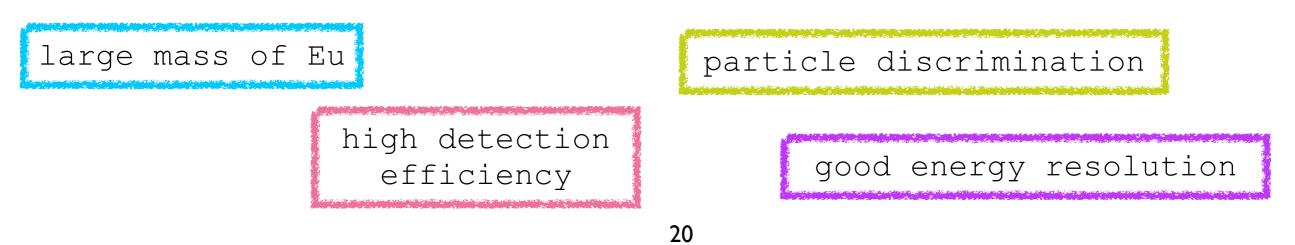
About 40% of the crystal mass is made of Eu: -> given <sup>151</sup>Eu isotopic abundance -> about 1.5 g of the crystal is made of <sup>151</sup>Eu

α-decay of <sup>151</sup>Eu never observed, just an indication in: Nucl. Phys. A 789 (2007) 15-29

#### Abstract

The indication for the  $\alpha$  decay of <sup>151</sup>Eu (Q $\alpha$  = 1.964 MeV) with the half-life T<sub>1/2</sub>  $\alpha$  = 5<sub>-3</sub><sup>+11</sup> × 10<sup>18</sup> yr has been observed for the first time with the help of a low background CaF<sub>2</sub>(Eu) crystal scintillator (mass of 370 g) in measurement at the Gran Sasso National Laboratories of the INFN during 7426 h. In a conservative approach the lower limit on the half-life of <sup>151</sup>Eu has been established as T<sup> $\alpha$ </sup> ≥ 1.7 × 10<sup>18</sup> yr at 68% C.L.

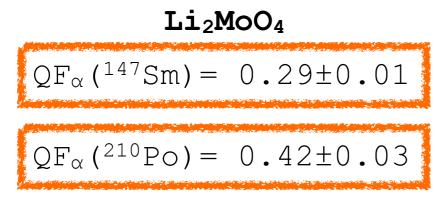
The discovery of this decay is not far away... ... Li<sub>6</sub>Eu(BO<sub>3</sub>)<sub>3</sub> scintillating bolometer seems to be the perfect tool



#### Conclusions

- Li-scintillating bolometers are a suitable tool for low background physics from DBD to solar axions

- the double read-out brings an abrupt reduction of the background in the ROI



- \* high radiopurity level
- \* good energy resolution

$$QF_{\alpha}(^{147}Sm) = 0.54\pm0.01$$

 $QF_{\alpha}(^{210}Po) = 0.84\pm0.05$ 

\* low radiopurity level
\* poor energy resolution

... bright future is ahead
but still some work is needed ...

