

# Radioactive contamination of $\text{CaWO}_4$ , $\text{ZnWO}_4$ , $\text{CdWO}_4$ , and $\text{Gd}_2\text{SiO}_5:\text{Ce}$ crystal scintillators

F.A. Danevich\*, A.Sh. Georgadze\*, V.V. Kobychiev\*, B.N. Kropivyansky\*, S.S. Nagorny\*, A.S. Nikolaiko\*, D.V. Poda\*, V.I. Tretyak\*, S.Yu. Zdesenko\*, Yu.G. Zdesenko<sup>1</sup>\*, P.G. Bizzeti<sup>†</sup>, T.F. Fazzini<sup>†</sup>, P.R. Maurenzig<sup>†</sup>, I.M. Solsky\*\*, V.B. Brudanin<sup>‡</sup> and F.T. Avignone III<sup>§</sup>

\**Institute for Nuclear Research, MSP 03680 Kiev, Ukraine*

<sup>†</sup>*Dipartimento di Fisica, Università di Firenze and INFN, 50019 Firenze, Italy*

\*\**Institute for Materials, 79031 Lviv, Ukraine*

<sup>‡</sup>*Joint Institute for Nuclear Research, 141980 Dubna, Russia*

<sup>§</sup>*University of South Carolina, Columbia, South Carolina 29208*

## Abstract.

Radioactive contaminations of  $\text{CaWO}_4$ ,  $\text{ZnWO}_4$ ,  $\text{CdWO}_4$ , and  $\text{Gd}_2\text{SiO}_5:\text{Ce}$  crystal scintillators were measured in the Solotvina Underground Laboratory. The Monte Carlo simulation of contributing radioactive sources, the time-amplitude and pulse-shape analyses of the experimental data were applied.  $\text{CaWO}_4$  and GSO scintillators are considerably polluted by uranium and thorium, while the total  $\alpha$  activity in  $\text{ZnWO}_4$  and  $\text{CdWO}_4$  scintillators does not exceed the level of a few mBq/kg. Particular radioactivity was observed in  $\text{CaWO}_4$  ( $\alpha$  active  $^{180}\text{W}$ ),  $\text{CdWO}_4$  ( $\beta$  active  $^{113}\text{Cd}$ ,  $2\nu 2\beta$  decay of  $^{116}\text{Cd}$ ,  $^{180}\text{W}$ ), and GSO ( $\alpha$  active  $^{152}\text{Gd}$ ) detectors. The radioactive contamination of the studied scintillation crystals is compared with that of commonly used detectors.

**Keywords:** Scintillation detectors; Radioactive contamination

**PACS:** PACS numbers: 29.40.Mc

## INTRODUCTION

There is class of experiments on searches for rare events (as for instance, low-level gamma-, beta-, alpha- spectrometry, double beta decay ( $2\beta$ ) and dark matter particles search, measurements of solar neutrino flux) which demand ultra-low background of detector. After the main sources of background (such as environmental radioactivity, cosmic rays, radioactivity of shield) are suppressed, internal radioactive contamination of detector begins to be the most important source of background.

Many experiments to search for rare  $\alpha$  [1, 2, 3] and  $\beta$  decay [4, 5, 9],  $2\beta$  decay [6, 8, 10, 11, 12, 13, 14, 15] and dark matter particles [16, 17, 10, 18] were performed using crystal scintillators. Some of scintillation crystals are also used as low-background scintillating bolometers [19, 20, 21, 22]. Here we analyze number of low-background measurements to estimate radioactive contamination of calcium, zinc, cadmium tungstate ( $\text{CaWO}_4$ ,  $\text{ZnWO}_4$ ,  $\text{CdWO}_4$ ), and cerium-doped gadolinium orthosilicate ( $\text{Gd}_2\text{SiO}_5:\text{Ce}$ , GSO) crystal scintillators as promising detectors for rare events experiments.

## EXPERIMENTAL METHODS

$\text{CaWO}_4$ ,  $\text{ZnWO}_4$ ,  $\text{CdWO}_4$ , and GSO scintillation crystals studied in the present work were grown by Czochralski method. The main properties of the scintillators are presented in Table 1. All these crystals are non-hygroscopic and chemically resistant.

---

<sup>1</sup> Deceased.

**TABLE 1.** Properties of CaWO<sub>4</sub>, ZnWO<sub>4</sub>, CdWO<sub>4</sub>, and GSO crystal scintillators

	CaWO <sub>4</sub>	ZnWO <sub>4</sub>	CdWO <sub>4</sub>	GSO
Density (g/cm <sup>3</sup> )	6.1	7.8	8.0	6.71
Melting point (° C)	1570 – 1650	1200	1325	2173
Hardness (Mohs)	4.5 – 5	4 – 4.5	4 – 4.5	5.7
Wavelength of emission maximum (nm)	420 – 440	480	480	440
Refractive index	1.94	2.1 – 2.2	2.2 – 2.3	1.85
Effective average decay time (μs)*	8	24	13	0.05
Photoelectron yield [% of NaI(Tl)]*	18%	13%	20%	20%

\* For  $\gamma$  rays at 20° C.

Radioactive contaminations of the crystals were measured in the Solotvina Underground Laboratory built in a salt mine 430 m underground ( $\approx 1000$  m w. e., cosmic muon flux  $1.7 \times 10^{-6}$  cm<sup>-2</sup> s<sup>-1</sup>, neutron flux  $\leq 2.7 \times 10^{-6}$  cm<sup>-2</sup> s<sup>-1</sup>, radon concentration in air  $< 30$  Bq m<sup>-3</sup>) [23].

To estimate radiopurity of cadmium tungstate scintillators, the data of the <sup>116</sup>Cd and <sup>106</sup>Cd double beta decay experiments [24, 8, 25, 12] were used. The radioactive contamination of GSO scintillator was determined on the basis of the experiment [26, 11].

The CaWO<sub>4</sub> (40 × 34 × 23 mm) was measured during 1734 h in the low background set-up. The CaWO<sub>4</sub> scintillator was viewed by the special low radioactive 5" PMT (EMI D724KFL) through the high pure quartz light-guide 10 cm in diameter and 33 cm long. The CdWO<sub>4</sub> crystal ( $\varnothing 59 \times 21$  mm) was measured during 37 h in the same set-up. The ZnWO<sub>4</sub> scintillation crystal ( $\varnothing 14 \times 4$  mm, 429 h) was viewed by the 3" PMT (FEU-137) through the high pure quartz light-guide 4.9 cm in diameter and 25 cm long. The detectors were surrounded by an passive shield made of teflon (3 – 5 cm), plexiglass (6 – 13 cm), high purity copper (thickness 3 – 6 cm), lead (15 cm) and polyethylene (8 cm). Two plastic scintillators (120 × 130 × 3 cm) were installed above the passive shield to generate a cosmic muons veto. For each event in the detector the amplitude of a signal and its arrival time were recorded. In addition, scintillation pulse shape of the CaWO<sub>4</sub> and CdWO<sub>4</sub> scintillators were digitized with a 20 MHz sampling frequency.

Pulse shape analysis was applied to CaWO<sub>4</sub> and CdWO<sub>4</sub> data as described in [27, 1, 28, 15]. Technique of time-amplitude analysis to recognize a presence of the short-living chains from <sup>232</sup>Th, <sup>235</sup>U and <sup>238</sup>U families is described in [24, 11]. The Monte Carlo simulation of contributing radioactive sources was performed with the help of GEANT3 and GEANT4 packages [29], and event generators DECAY0 and DECAY4 [30].

## RESULTS AND DISCUSSION

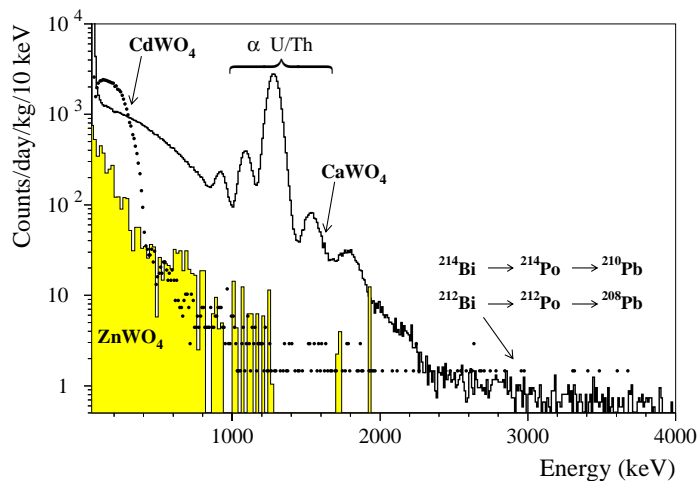
The energy spectrum measured with the ZnWO<sub>4</sub>, CaWO<sub>4</sub> and CdWO<sub>4</sub> crystal scintillators in the low background set-up is presented in Fig. 1 (the spectra are normalized by the measurement times and the detector masses).

The background of the ZnWO<sub>4</sub> detector is substantially lower than that of the CaWO<sub>4</sub> and is comparable with that of the CdWO<sub>4</sub> above  $\approx 0.5$  MeV. Note that below 0.5 MeV the counting rate of the ZnWO<sub>4</sub> detector is one order of magnitude lower than that of CdWO<sub>4</sub>. Obviously, it is due to presence of the  $\beta$  active <sup>113</sup>Cd isotope (natural abundance of  $\approx 12\%$ ) in the CdWO<sub>4</sub> crystals.

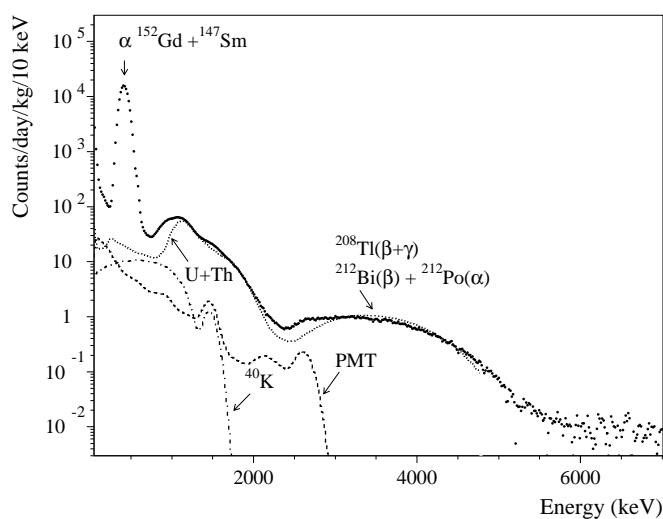
Activity of thorium and uranium  $\alpha$  active daughters in CdWO<sub>4</sub> and CaWO<sub>4</sub> crystals was determined with the help of pulse shape analysis. In addition, the pulse shape analysis allows to estimate activity of <sup>228</sup>Th and <sup>226</sup>Ra due to selection of double pulses associated with the fast sequences <sup>214</sup>Bi  $\rightarrow$  <sup>214</sup>Po ( $T_{1/2} = 164$  μs)  $\rightarrow$  <sup>210</sup>Pb and <sup>212</sup>Bi  $\rightarrow$  <sup>212</sup>Po ( $T_{1/2} = 0.3$  μs)  $\rightarrow$  <sup>208</sup>Pb from <sup>235</sup>U chain.

More precise data on radioactive impurities associated with the daughters of <sup>232</sup>Th, <sup>235</sup>U and <sup>238</sup>U were obtained with the help of the time-amplitude analysis. For example, the fast sequence of two  $\alpha$  decays from the <sup>232</sup>Th family was searched for: <sup>220</sup>Rn ( $Q_{\alpha} = 6.41$  MeV,  $T_{1/2} = 55.6$  s)  $\rightarrow$  <sup>216</sup>Po ( $Q_{\alpha} = 6.91$  MeV,  $T_{1/2} = 0.145$  s)  $\rightarrow$  <sup>212</sup>Pb (which is in equilibrium with <sup>228</sup>Th). In the same way, the activity of <sup>226</sup>Ra (<sup>238</sup>U family) in the scintillators was measured by selection of the fast sequence <sup>214</sup>Bi ( $Q_{\beta} = 3.27$  MeV,  $T_{1/2} = 19.9$  m)  $\rightarrow$  <sup>214</sup>Po ( $Q_{\alpha} = 7.83$  MeV,  $T_{1/2} = 164$  μs)  $\rightarrow$  <sup>210</sup>Pb. The presence of <sup>227</sup>Ac (<sup>235</sup>U family) was estimated by time-amplitude analysis of <sup>219</sup>Rn ( $Q_{\alpha} = 6.95$  MeV,  $T_{1/2} = 3.96$  s)  $\rightarrow$  <sup>215</sup>Po ( $Q_{\alpha} = 7.53$  MeV,  $T_{1/2} = 1.78$  ms)  $\rightarrow$  <sup>211</sup>Pb from <sup>235</sup>U chain.

The background spectrum measured with the GSO crystal ( $\varnothing 47 \times 57$  mm) during 13949 h [11] is depicted in Fig.



**FIGURE 1.** Energy spectra of ZnWO<sub>4</sub> (4.5 g, 429 h), CaWO<sub>4</sub> (189 g, 1734 h), and CdWO<sub>4</sub> (448 g, 37 h) scintillation crystals measured in the low background set-up



**FIGURE 2.** The background spectrum of the GSO crystal for 370.5 yr kg of exposure and the model of background.

2, where the following peculiarities exist: the clear peak at the energy 0.42 MeV, the comparatively wide peak at the energy around 1.05 MeV and two broad distributions dropped down at the energies 2.4 and 5.5 MeV. The first peak is attributed to  $\alpha$  particles of  $^{152}\text{Gd}$  ( $T_{1/2} = 1.08 \times 10^{14}$  yr;  $E_{\alpha} = 2140$  keV; abundance  $\delta = 0.20\%$ ) and  $^{147}\text{Sm}$  ( $T_{1/2} = 1.06 \times 10^{11}$  yr;  $E_{\alpha} = 2233$  keV;  $\delta = 15\%$ ; samarium can be present as impurity of the GSO crystal at the level of  $\approx 8$  ppm). The peak near 1050 keV as well as the broad distribution up to the energy 2.4 MeV is mainly due to the radioactive contamination of the crystal by the nuclides from the  $^{232}\text{Th}$ ,  $^{235}\text{U}$  and  $^{238}\text{U}$  families. The distribution up to the energy 5.5 MeV is caused by decays of  $^{232}\text{Th}$  daughter isotopes: a)  $\beta$  decay of  $^{208}\text{Tl}$  ( $Q_{\beta} = 5.00$  MeV); b)  $\beta$  decay of  $^{212}\text{Bi}$  ( $Q_{\beta} = 2.25$  MeV) followed by fast  $\alpha$  decay of its daughter  $^{212}\text{Po}$  ( $T_{1/2} = 0.3 \mu\text{s}$ ;  $E_{\alpha} = 8.78$  MeV or  $\approx 2.7$  MeV in  $\beta$  scale). The estimation of radioactive contamination of the GSO scintillator was performed by using the time-amplitude analysis and by analysis of the energy spectrum shape. The model of background was built with the help of Monte Carlo simulation. It should be stressed that no surface concentration of radionuclides was observed for the GSO crystal.

The summary of the measured radioactive contamination of the CaWO<sub>4</sub>, ZnWO<sub>4</sub>, CdWO<sub>4</sub> and GSO crystal scintillators (or limits on their activities) is given in Table 2.

As one can see from Table 2, the secular equilibrium in uranium chain is substantially broken in CaWO<sub>4</sub> crystal.

**TABLE 2.** Measured radioactive contaminations in ZnWO<sub>4</sub>, CaWO<sub>4</sub>, CdWO<sub>4</sub>, and GSO crystal scintillators (mBq/kg).

Chain	Source	CaWO <sub>4</sub> *	ZnWO <sub>4</sub> †	CdWO <sub>4</sub> **	GSO‡
<sup>232</sup> Th	<sup>232</sup> Th	0.69(10)	≤ 3.3	0.053(5)	≤ 6.5
	<sup>228</sup> Ra				≤ 9
	<sup>228</sup> Th	0.6(2)	≤ 0.2	≤ 0.004 – 0.039(2)	2.287(13)
<sup>235</sup> U	<sup>227</sup> Ac	1.6(3)	≤ 0.2	0.0014(9)	0.948(9)
<sup>238</sup> U	<sup>238</sup> U	14.0(5)	≤ 3.2	≤ 0.6	≤ 2
	<sup>230</sup> Th		≤ 4.5	≤ 0.5	≤ 9
	<sup>226</sup> Ra	5.6(5)	≤ 0.4	≤ 0.004	0.271(4)
	<sup>210</sup> Pb	≤ 430	≤ 1	≤ 0.4	≤ 0.8
	<sup>210</sup> Po	291(5)			
	<sup>40</sup> K	≤ 12	≤ 12	0.3(1)	≤ 14
	<sup>90</sup> Sr	≤ 70	≤ 1.2	≤ 0.2	
	<sup>113</sup> Cd			580(20) §	
	<sup>113m</sup> Cd			1-30	
	<sup>116</sup> Cd			1.0(1) ¶	
<sup>137</sup> Cs	≤ 0.8	≤ 20	≤ 0.3 – 0.43(6)		
<sup>138</sup> La				≤ 55	
<sup>147</sup> Sm	0.49(4)	≤ 1.8	≤ 0.04	700-1100	
<sup>152</sup> Gd			≤ 0.04	1200	
<sup>180</sup> W	0.05(2)		0.04(2)		

\* [28, 31]

† [15]

\*\* [25, 32, 8, 12, 1]

‡ [26, 11]

§ In crystals produced from natural cadmium

¶ In crystals produced from Cd enriched in <sup>116</sup>Cd

The equilibrium in thorium family is broken in CdWO<sub>4</sub> scintillators too. More likely it is a rule for different materials. On the other hand, measurement of particular sub-chains with comparable precision is quite complicated problem. For that reason radioactive contamination by uranium and thorium is more properly characterized by the value of the total  $\alpha$  activity inside a detector.

Particular radioactivity, i.e. associated with elemental composition of a detector, was observed in CaWO<sub>4</sub> ( $\alpha$  active <sup>180</sup>W), CdWO<sub>4</sub> ( $\beta$  active <sup>113</sup>Cd,  $2\nu 2\beta$  decay of <sup>116</sup>Cd, <sup>180</sup>W), and GSO ( $\alpha$  active <sup>152</sup>Gd) detectors.

Let us mention the effect of U/Th traces concentration on surface layer of the CdWO<sub>4</sub> crystal grown from enriched <sup>116</sup>Cd. Contribution from U/Th  $\alpha$  active daughters was decreased in  $\approx 50$  times after removing of surface layer ( $\approx 1$  mm) of the crystal [24].

In Table 3 the radioactive contamination of the studied scintillation crystals is compared with that of commonly used detectors.

## CONCLUSIONS

We present the radio-purities of CaWO<sub>4</sub>, ZnWO<sub>4</sub>, CdWO<sub>4</sub> and GSO crystals and compare them with other detectors. CaWO<sub>4</sub> and GSO scintillators are considerably polluted by uranium and thorium (particularly, CaWO<sub>4</sub> by <sup>210</sup>Po at the level of 0.3 Bq/kg, GSO by U/Th at the level of 0.04 Bq/kg). Radiopurity of ZnWO<sub>4</sub> and CdWO<sub>4</sub> crystals is significantly better. The total alpha activity of U/Th daughters in ZnWO<sub>4</sub> and CdWO<sub>4</sub> scintillators does not exceed the level of a few mBq/kg. The contaminations of CdWO<sub>4</sub>, enriched in <sup>116</sup>Cd, by <sup>232</sup>Th and <sup>226</sup>Ra are at the level of 53  $\mu$ Bq/kg and 4  $\mu$ Bq/kg, respectively. Particular, i.e. associated with elemental composition of the detectors, radioactivity was observed in CaWO<sub>4</sub> ( $\alpha$  active <sup>180</sup>W), CdWO<sub>4</sub> ( $\beta$  active <sup>113</sup>Cd and <sup>113m</sup>Cd,  $2\nu 2\beta$  decay of <sup>116</sup>Cd,

**TABLE 3.** Radioactive contamination of different detectors (mBq/kg).

Detector	Total $\alpha$ activity U + Th	$^{228}\text{Th}$	$^{226}\text{Ra}$	$^{40}\text{K}$	Particular contamination	Reference
CaWO <sub>4</sub>	20–400	0.6	5.6	≤ 12		[28, 31, 22]
ZnWO <sub>4</sub>	≤ 20	≤ 0.2	≤ 0.4	≤ 12		[15]
CdWO <sub>4</sub>	≤ 0.7–2.3(3)	≤ 0.003	≤ 0.004	0.3	580 ( $^{113}\text{Cd}$ )	[25, 32, 8, 12, 1]
GSO	40	2.3–107	0.3	≤ 14	1200 ( $^{152}\text{Gd}$ )	[26, 11, 33]
NaI(Tl)	≤ 0.2	0.0004–0.014	0.005–0.39	1.5		[34, 35, 36]
CsI(Tl)		≤ 0.3	≤ 0.3	≤ 20		[37, 38]
BGO	≤ 0.4	≤ 1.2			(0.5–3) × 10 <sup>3</sup> ( $^{207}\text{Bi}$ )	[39, 2]
CaF <sub>2</sub> (Eu)		0.04–40	0.05–75	5		[14, 10]
CeF <sub>3</sub>	3400	1100	≤ 60	≤ 330		[13]
BaF <sub>2</sub>		400	1400			[40]
Liquid scintillator	10 <sup>-6</sup>	1.2 × 10 <sup>-6</sup>	6.3 × 10 <sup>-6</sup>		0.3 ( $^{14}\text{C}$ )	[41, 42]
HPGe		≤ 2 × 10 <sup>-5</sup>	≤ 2 × 10 <sup>-5</sup>			[43, 44]

$^{180}\text{W}$ ), and GSO ( $\alpha$  active  $^{152}\text{Gd}$ ) detectors. It was found that secular equilibrium in uranium chain is substantially broken in CaWO<sub>4</sub> crystals. Uranium and thorium traces were concentrated mainly in the surface layer of the CdWO<sub>4</sub> crystals enriched by  $^{116}\text{Cd}$ .

## REFERENCES

1. F. A. Danevich, A. Sh. Georgadze, V. V. Kobychhev, S. S. Nagorny et al., *Phys. Rev. C*, **67**, 014310, 8 p (2003).
2. P. de Marcillac, *Nature*, **422**, 876–878 (2003).
3. C. Cozzini, G. Angloher, C. Bucci, F. von Feilitzsch et al., *Phys. Rev. C* **70**, 064606, 6 p (2004).
4. A. Alessandrello, C. Brofferio, D. V. Camin, O. Cremonesi et al., *Nucl. Phys. B (Proc. Suppl.)*, **35**, 394–396 (1994).
5. F. A. Danevich, A. Sh. Georgadze, V. V. Kobychhev, B. N. Kropivnyansky et al., *Phys. At. Nucl.*, **59**, 1–5 (1996).
6. F. A. Danevich, Yu. G. Zdesenko, A. S. Nikolaiko, S. F. Burachas et al., *Instr. Exp. R.*, **32**, 1059–1064 (1989).
7. F. A. Danevich, Yu. G. Zdesenko, A. S. Nikolaiko, V. I. Tretyak, *JETP Lett.*, **49**, 476–479 (1989).
8. F. A. Danevich, A. Sh. Georgadze, J. Hellmig, M. Hirsch et al., *Z. Phys. A*, **355**, 433–437 (1996).
9. C. M. Cattadori, M. De Deo, M. Laubenstein, L. Pandola, V. I. Tretyak, *Nucl. Phys. A*, **748** 333–347 (2005).
10. P. Belli, R. Bernabei, C. J. Dai, F. Grianti et al., *Nucl. Phys. B*, **563**, 97–106 (1999).
11. F. A. Danevich, V. V. Kobychhev, O. A. Ponkratenko, V. I. Tretyak, Yu. G. Zdesenko, *Nucl. Phys. A*, **694**, 375–391 (2001).
12. F. A. Danevich, A. Sh. Georgadze, V. V. Kobychhev, B. N. Kropivnyansky et al., *Phys. Rev. C*, **68**, 035501, 12 p (2003).
13. P. Belli, R. Bernabei, R. Cerulli, C. J. Dai et al., *Nucl. Instr. Meth. A*, **498**, 352–361 (2003).
14. I. Ogawa, R. Hazama, H. Miyawaki, S. Shiomi et al., *Nucl. Phys. A*, **730**, 215–223 (2004).
15. F. A. Danevich, V. V. Kobychhev, S. S. Nagorny, D. V. Poda et al., nucl-ex/0409014; *Nucl. Instr. Meth. A*, (2005) in press.
16. C. Bacci, P. Belli, R. Bernabei, C. J. Dai et al., *Astropart. Phys.*, **2**, 13–19 (1994).  
R. Bernabei, P. Belli, V. Landoni, F. Montecchia et al., *Phys. Lett. B*, **389**, 757–766 (1996).  
R. Bernabei, P. Belli, F. Montecchia, W. Di Nicolantonio et al., *Phys. Lett. B*, **424**, 195–201 (1998).  
R. Bernabei, P. Belli, F. Montecchia, W. Di Nicolantonio et al., *Phys. Lett. B*, **450**, 448–455 (1999).  
R. Bernabei, P. Belli, R. Cerulli, F. Montecchia et al., *Phys. Lett. B*, **480**, 23–31 (1998).  
R. Bernabei, M. Amato, P. Belli, R. Cerulli et al., *Phys. Lett. B*, **509**, 197–203 (2001).  
R. Bernabei, P. Belli, R. Cerulli, F. Montecchia et al., *Eur. Phys. J. C*, **23**, 61–64 (2002).  
R. Bernabei, P. Belli, F. Cappella, R. Cerulli et al., *Riv. Nuovo Cim.*, **26**, 1–73 (2003).
17. C. Bacci, P. Belli, R. Bernabei, C. J. Dai et al., *Astropart. Phys.*, **2**, 117–125 (1994).
18. I. Ogawa, T. Kishimoto, R. Hazama, S. Ajimura et al., *Nucl. Phys. A*, **663&664**, 869c–872c (2000).
19. A. Alessandrello, D. V. Camin, O. Cremonesi, E. Fiorini et al., *Nucl. Phys. B (Proc. Suppl.)*, **28A**, 233–235 (1992).  
C. Arnaboldi, C. Brofferio, C. Bucci, S. Capelli et al., *Phys. Lett. B*, **557**, 167–175 (2003).
20. C. Bobin, I. Berkes, J. P. Hadjout, N. Coron et al., *Nucl. Instr. Meth. A*, **386**, 453–457 (1997).
21. M. Bravin, M. Bruckmayer, C. Bucci, S. Cooper et al., *Astropart. Phys.* **12**, 107–114 (1999).
22. S. Cebrían N. Coron, G. Dambier, P. de Marcillac et al., *Phys. Lett. B*, **563**, 48–52 (2003).
23. Yu. G. Zdesenko, B. N. Kropivnyansky, V. N. Kuts, A. S. Nikolaiko et al., in *Proc. 2nd Int. Symp. Underground Physics, Baksan Valley, USSR, August 17–19, 1987*, edited by G. V. Domogatsky, Nauka, Moscow, 1988, pp. 291–295.
24. F. A. Danevich, A. Sh. Georgadze, V. V. Kobychhev, B. N. Kropivnyansky et al., *Phys. Lett. B* **344**, 72–78 (1995).
25. A. Sh. Georgadze, F. A. Danevich, Yu. G. Zdesenko, V. V. Kobychhev et al., *Prib. Tekh. Exp.* **2**, 45–51 (1996) [*Instrum. Exp. Tech.* **39**, 191–198 (1996)].

26. S. F. Burachas, F. A. Danevich, Yu. G. Zdesenko, V. V. Kobychyev et al., *Phys. Atom. Nucl.*, **58**, 153–157 (1995).
27. T. Fazzini, P. G. Bizzeti, P. R. Maurenzig, C. Stramaccioni et al., *Nucl. Instr. Meth. A*, **410**, 213–219 (1998).
28. Yu. G. Zdesenko, F. T. Avignone III, V. B. Brudanin, F. A. Danevich et al., *Nucl. Instr. Meth. A*, **538**, 657–667 (2005).
29. CERN Program Library Long Write-up W5013, CERN, 1994;  
S. Agostinelli, J. Allison, K. Amako, J. Apostolakis et al., (GEANT4 Collaboration), *Nucl. Instr. Meth. A*, **506**, 250–303 (2003), URL <http://geant4.web.cern.ch/geant4/>.
30. O. A. Ponkratenko, V. I. Tretyak, Yu. G. Zdesenko, *Yad. Fis.*, **63**, 1355–1360 (2000) [*Phys. At. Nucl.*, **63**, 1282–1287 (2000)].
31. Yu. G. Zdesenko, F. T. Avignone III, V. B. Brudanin, F. A. Danevich et al., *Astropart. Phys.*, **23**, 249–263 (2005).
32. S. Ph. Burachas, F. A. Danevich, A. Sh. Georgadze, H. V. Klapdor-Kleingrothaus et al., *Nucl. Instr. Meth. A*, **369**, 164–168 (1996).
33. S. C. Wang, H. T. Wong and M. Fujiwara, *Nucl. Instr. Meth. A*, **479**, 498–510 (2002).
34. R. Bernabei, P. Belli, F. Montecchia, W. Di Nicolantonio et al., *Il Nuovo Cim. A*, **112**, 545–564 (1999).
35. J. C. Barton, J. A. Edgington, *Nucl. Instr. Meth. A*, **443**, 277–286 (2000).
36. S. Umehara, K. Fushimi, N. Koori, N. Kudomi et al., *Nucl. Instr. Meth. A*, **490**, 271–275 (2002).
37. A. Baranov, V. Bashkirov, G. Bondarenko, A. Strifutkin, *Preprint MEPI 040-89*, Moscow Engineering Physics Institute, Moscow, 1989.
38. H. T. Wong, J. Li, C. Y. Chang, C. C. Chang et al., *Astropart. Phys.*, **14**, 141–152 (2000).
39. A. Balysh, A. Gurov, A. Demehin, I. Kondratenko et al., *Prib. Tekh. Exp.*, **1**, 118–122 (1993).
40. R. Cerulli, P. Belli, R. Bernabei, F. Capella et al., *Nucl. Instr. Meth. A*, **525**, 535–543 (2004).
41. Borexino Collaboration, *Borexino Status Report*, LNGS, September 30, 2004 (unpublished).
42. G. Alimonti, G. Angloher, C. Arpesella, M. Balata et al., *Phys. Lett. B*, **422**, 349–358 (1998).
43. H. V. Klapdor-Kleingrothaus, L. Baudis, A. Dietz, G. Heusser et al., *Nucl. Instr. Meth. A*, **481**, 149–159 (2002).
44. C. Dörr, H. V. Klapdor-Kleingrothaus, *Nucl. Instr. Meth. A*, **513**, 596–621 (2003).